The GEOS RR Lyr database and survey A professional-amateur project

Jean-François Le Borgne

GEOS
Groupe Européen d'Observation Stellaire
IRAP

Institut de Recherche en Astrophysique et Planétologie Observatoire Midi-Pyrénées, Université Paul Sabatier Toulouse, France



GEOS: http://geos.upv.es/

- Created in 1974, already European and pro-am.
- Promote research among amateur astronomers in Europe
- The basic idea is that amateurs should themselves extract scientific information from their observations and publish their results.
- Recent initiative: creation of a RR Lyr star maxima database.
- During the past years the study of RR Lyr stars has become the main field of interest of GEOS but other variable stars are also studied: eclipsing binaries, δ Scuti stars, Cepheids ... essentially "fast" varying stars

References:

Hambsch, F.-J.; Le Borgne, J.-F.; Poretti, E.; GEOS association, 2012, The Journal of the American Association of Variable Star Observers, vol. 40, no. 1, p. 177

Boninsegna, R.; Vandenbroere, J.; Le Borgne, J. F.; GEOS Team, 2002, ASP Conference Proceedings, Vol. 259. 166

GEOS RR Lyr Database

A tool for the study of RR Lyr stars

http://rr-lyr.irap.omp.eu/dbrr/

Allow to follow period variations since star's discovery, some times over 100 years ago.

Help to preparation of observations of RR Lyr stars,

GEOS RR Lyr Survey

Add significantly more maxima timings of the brightest RR Lyr stars essentially using robotic telescopes.

Study fainter understudied stars to refine their period and find new stars which exhibit Blazhko effect.

Characterize the Blazhko effect, and other long term variations of RR Lyr stars.

GEOS RR Lyr Database

Aim of the database: Maintaining an up-to-date list of observed maxima of RR Lyr stars.

Starting idea in 1999: updating the data published in the book "RR Lyr Stars" written by Vladimir P. Tsesevich in the years 1960's (1966, 1969 for the English translation).

And make it available on the web,

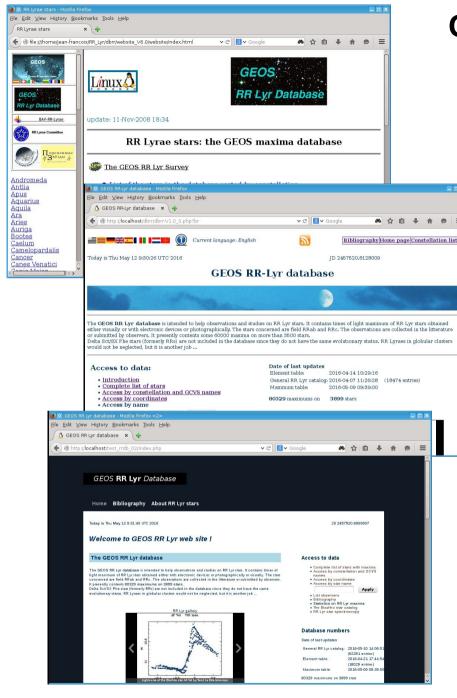
First, collecting published historical times of maximum made with the efficient help of Anton Paschke, Massimiliano Martignoni and Francesco Acerbi in the early years of the project.

We make a survey of recent publications periodically

A couple of days ago: 81714 maxima on 3904 stars

To access data: a **web interface** to the database list of maxima, O-C curves, basic data of the stars

The web site is hosted by IRAP.



GEOS RR Lyr Database: 2000-2016 3 versions

V1:

Interface coded in html, static Maximum data in ascii files

V2:

Interface coded in html + php
Web pages created dynamically
Maximum data in database mysql
Indexed on gcvs star names

V3:

Interface coded in html + php + css
Web pages created dynamically
Maximum data in database mysql
Indexed on unique star identification
Uses alias names

Database tables

- Catalog: coordinates, type, magnitudes
- Alias names
- Elements
- Maximums
- References
- Comments

Web interface

Access to data:

- Access by constellation and GCVS names
- Access by star name
- Complete list of stars with maxima
- Access by coordinates / Access to catalog
 - Center of field or field limits
 - Center on given star

Access to data

- Complete list of stars with maxima.
- Access by constellation and GCVS names
- · Access by coordinates
- Access by star name

Apply

- List observers
- Bibliography
- · Statistics on RR Lyr maxima
- The Blazhko star catalog
- RR Lyr star spectroscopy

Database numbers

Date of last updates

General RR Lyr catalog: 2016-05-10 14:06:51

(62281 entries)

Element table: 2016-04-21 17:44:54

(18029 entries)

Maximum table: 2016-05-13 13:23:27

80337 maximums on 3899 stars

Making of the RR Lyr catalog

Access to GCVS and VSX catalogs

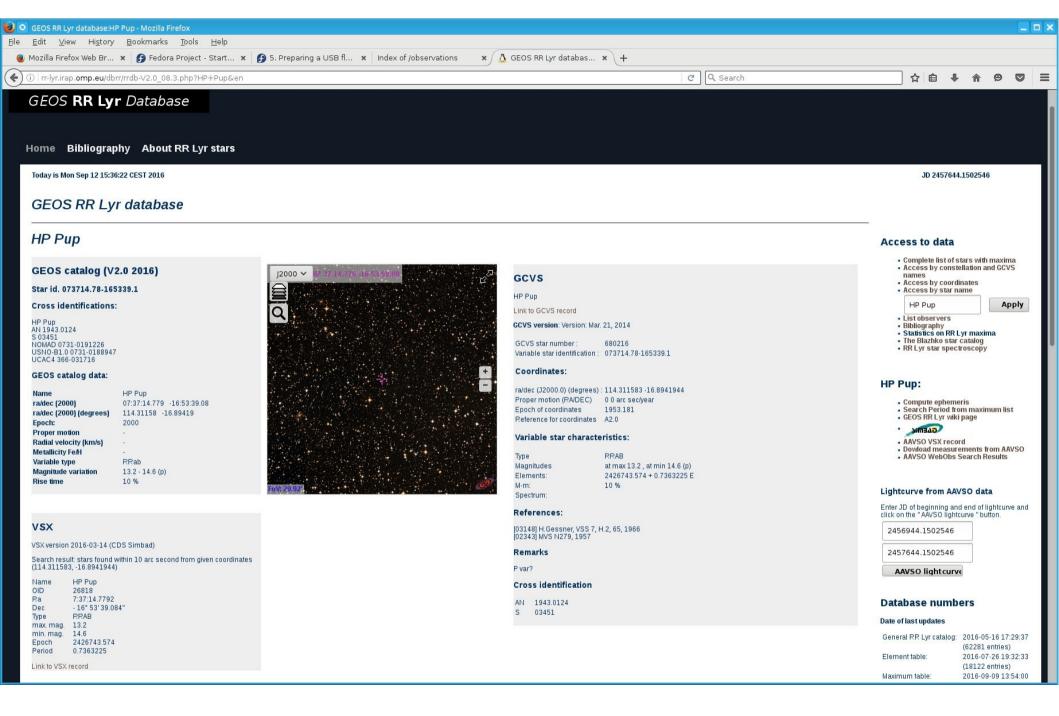
VSX catalog is updated every Monday at CDS GCVS is updated irregularly at scale of several

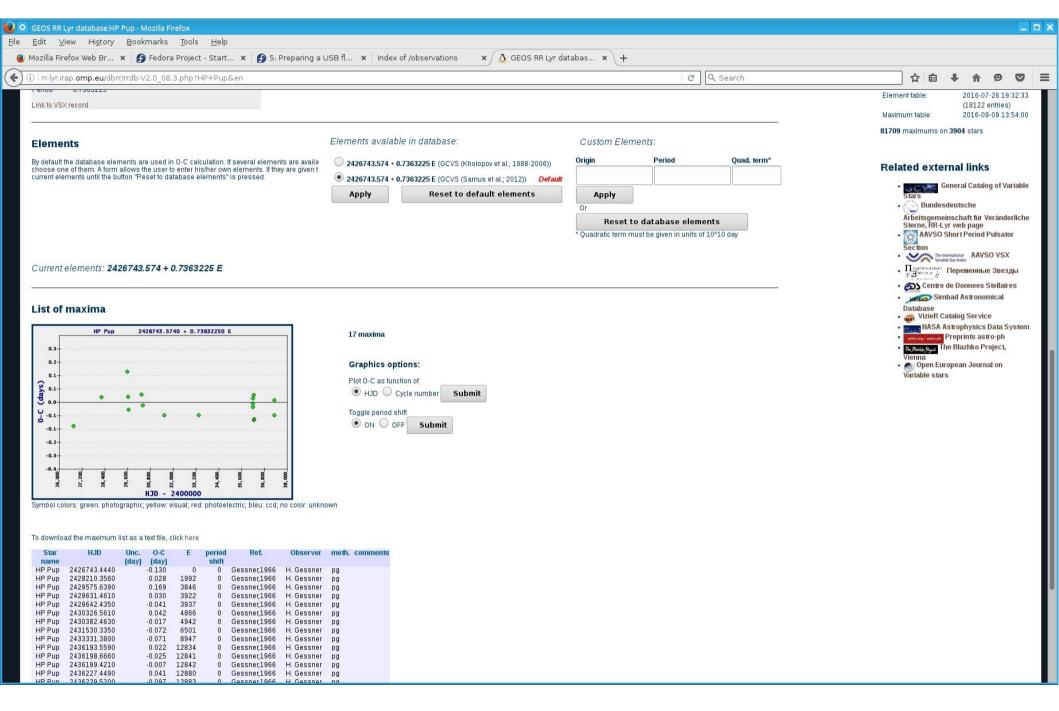
months

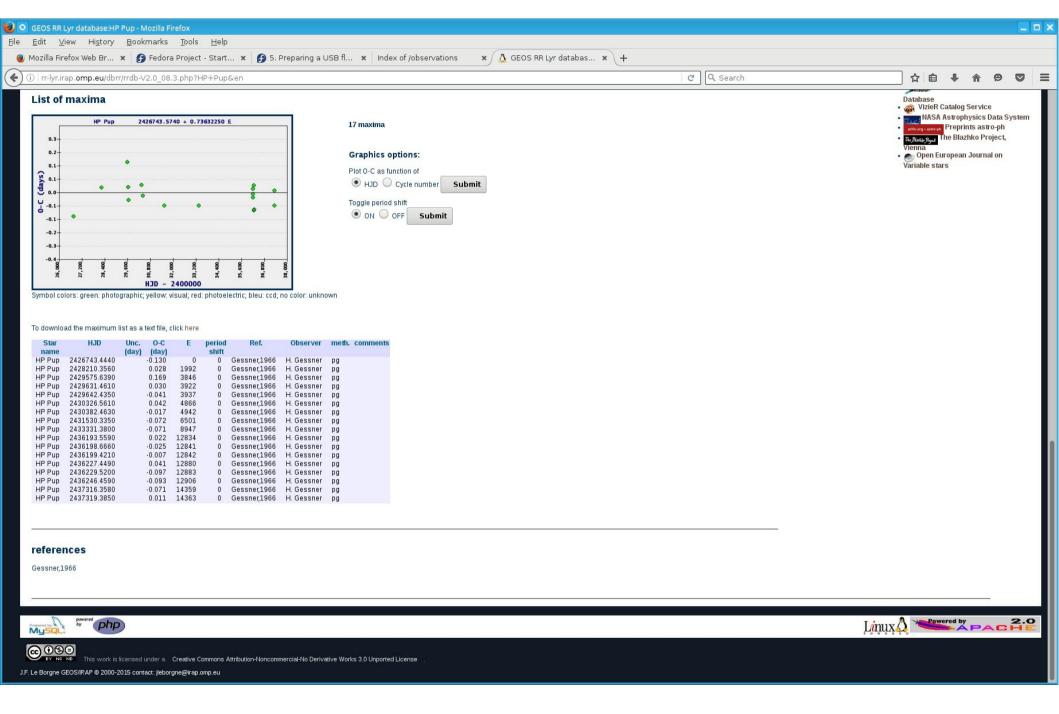


Index of ftp://cdsarc.u-strasbg.fr/pub/cats/B/vsx/









Access to data and functions

Access to data

- Complete list of stars with maxima
- Access by constellation and GCVS names
- Access by coordinates
- · Access by star name

SW And



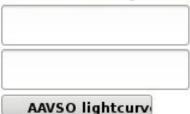
- List observers
- Bibliography
- Statistics on RR Lyr maxima
- The Blazhko star catalog
- RR Lyr star spectroscopy

HP Pup:

- Compute ephemeris
- Search Period from maximum list
- GEOS RR Lyr wiki page
- · MINISAD
- AAVSO VSX record
- Dowload measurements from AAVSO
- · AAVSO WebObs Search Results

Lightcurve from AAVSO data

Enter JD of beginning and end of lightcurve and click on the " AAVSO lightcurve " button.



Database numbers

Date of last updates

General RR Lyr catalog: 2016-05-16 17:29:37

(62281 entries)

Element table: 2016-07-26 19:32:33

(18122 entries)

Maximum table: 2016-09-09 13:54:00

81709 maximums on 3904 stars

Related external links

 General Catalog of Variable Stars

Bundesdeutsche
 Arbeitsgemeinschaft für
 Veränderliche Sterne, RR-Lyr web
 page

AAVSO Short Period Pulsator

Section

The International AAVSO VSX

• $\frac{\prod_{\text{временяне}}}{\sqrt[3]{3}}$ Переменные Звезды

• 🙉 Centre de Donnees Stellaires

• Simbad Astronomical

Database

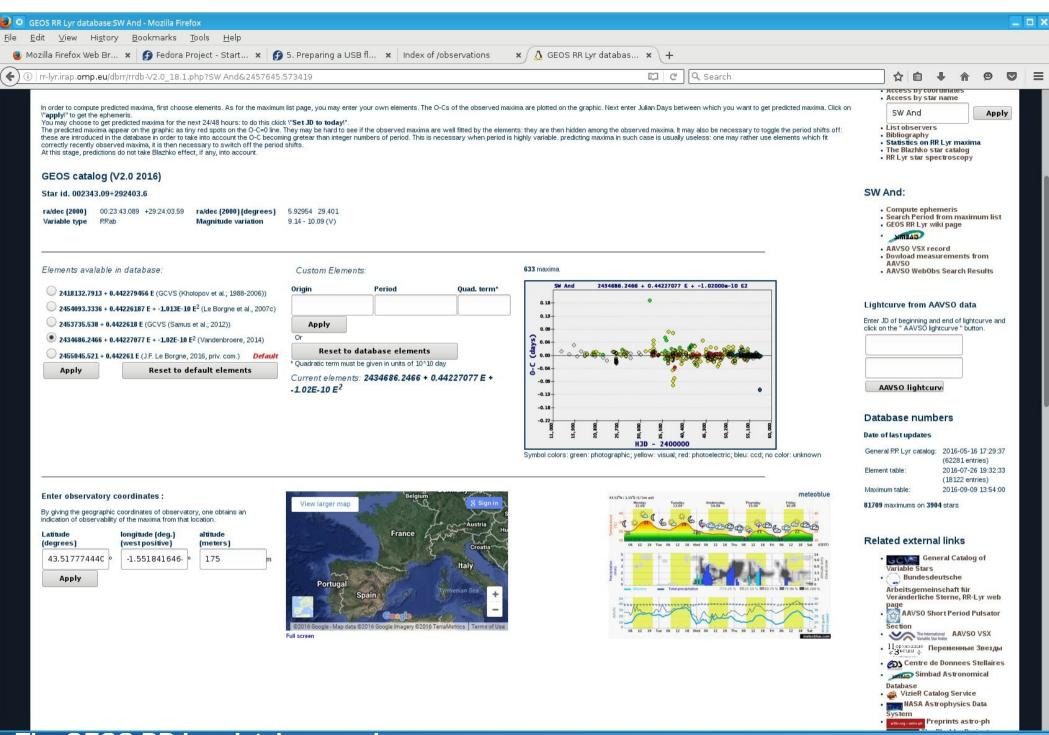
WizieR Catalog Service

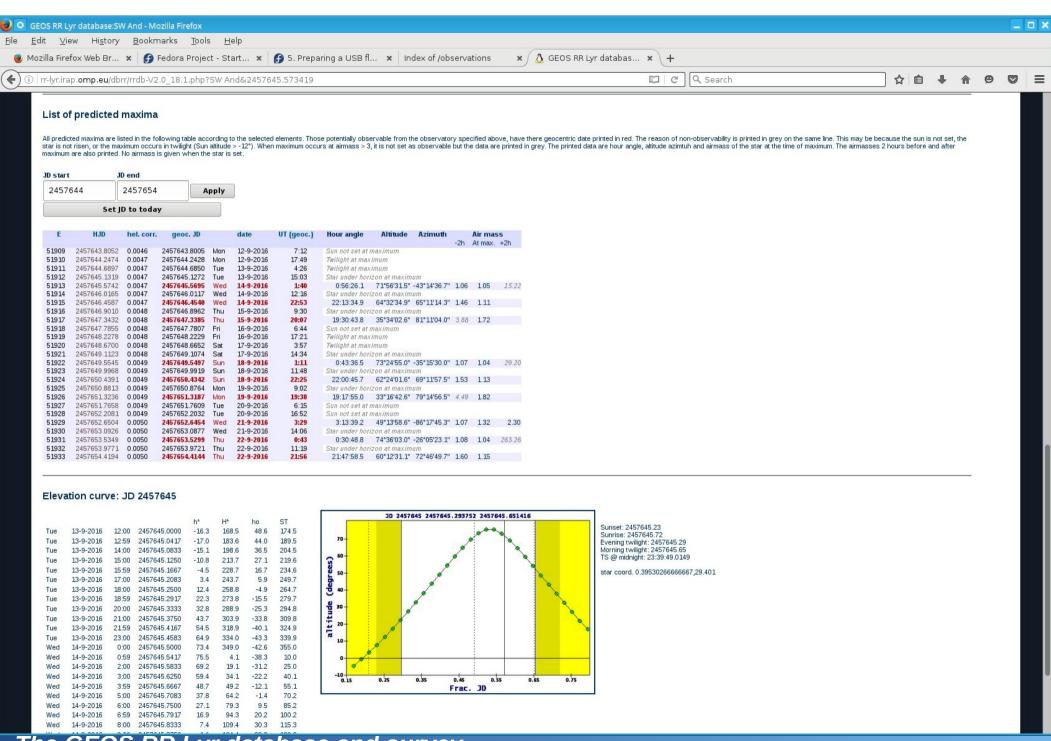
NASA Astrophysics Data System

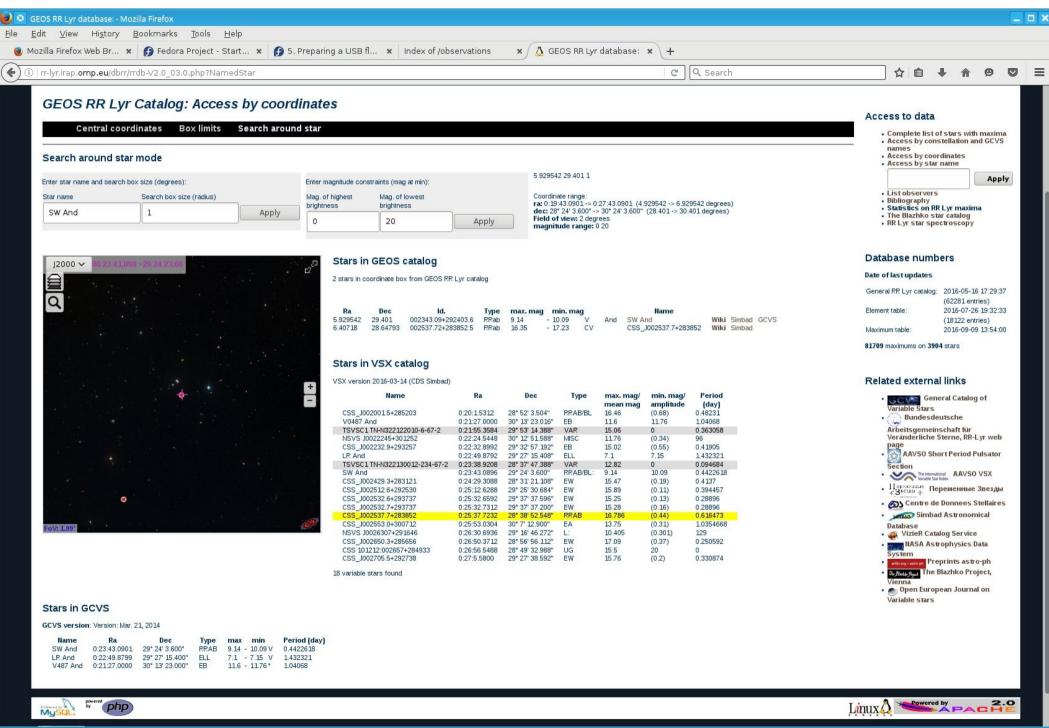
arXiv.org > astro-ph Preprints astro-ph

The Blazhko Project, Vienna

 Open European Journal on Variable stars







GEOS RR Lyr Survey

Started 2004

Robots: routine observations of maximums of bright RRab stars (magnitude at minimum <~13).

Aim: survey light curve variations at large time scales (>~10 years)

Human observers: Study of fainter understudied stars (magnitude at minimum range ~13-15).

Aim: refine or find period and find possible new Blazhko effects,

Human observers: Followup of RR Lyr itself with small dedicated instruments

Robotic Telescopes Tarot « Télescope à Action Rapide pour les Objets Transitoires »

PIs: M. Boër(OCA) and A. Klotz (IRAP)

Dedicated to followup of gamma ray bursts And observation of earth satellites for CNES. Mirror diameter 25 cm Robotic and fast CCD cameras 2000x2000, field 2°x2°

RR Lyr star survey is one of the additional programs

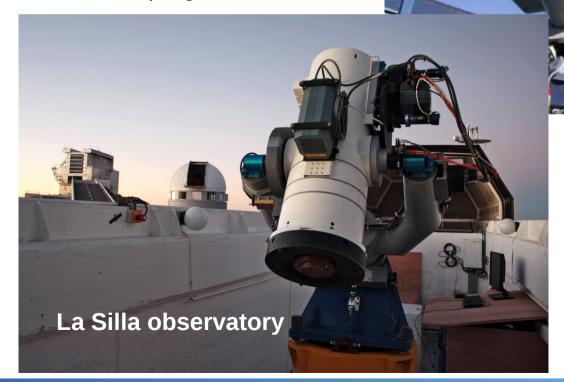
(~40% observing time)

Calern, France: since 2004 8557 maximums 294 stars

La Silla, Chile: since 2006 7364 maximums 360 stars

Magnitudes: 9 to 13

Automatic data reduction: bias, flat fields, photometry

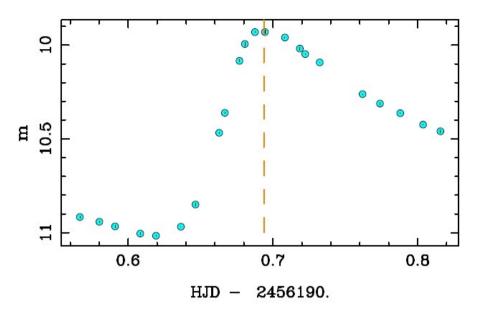


Calern observatory

Routine followup of bright RR Lyr stars of type RRab

6 to 8 maxima scheduled on each telescope every night

Precision of ToMs: ~0.002 days (3mn)

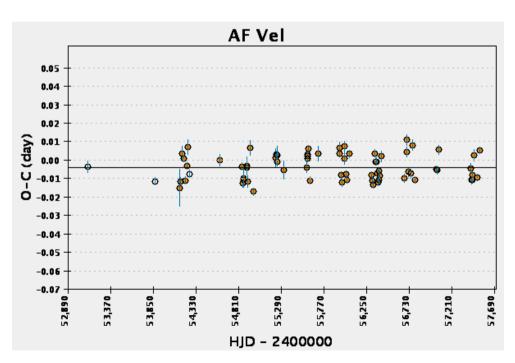


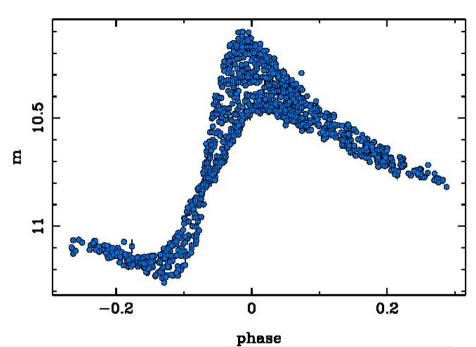
Thursday 3 March 2016

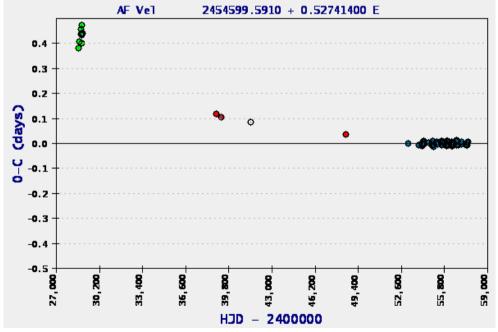
JD	Star	N*	observation time (JD)		observation time (UT)		Maximum HJD			
			from	to	from	to	Predicted	Observed	O-C (day)	Observer/Telescope
0455454	D0.15-	4.0	0.400	0.500	84.6					m
2457451		16	0.409	0.508	21.8	0.2				Tarot - Calern
2457451	RR Leo	26	0.312	0.517	19.5	0.4				Tarot - Calern
2457451	EX UMa	20	0.315	0.434	19.6	22.4	2457451.343	2457451.348± 0.004	0.005	Tarot - Calern
2457451	IM Leo	36	0.396	0.628	21.5	3.1	2457451.518	2457451.519± 0.002	0.001	Tarot - Calern
2457451	RV UMa	28	0.560	0.706	1.4	4.9	2457451.619	2457451.622± 0.002	0.002	Tarot - Calern
2457451	RX Leo	48	0.466	0.707	23.2	5.0	2457451.570	2457451.574± 0.003	0.004	Tarot - Calern
2457451	TW Boo	38	0.578	0.698	1.9	4.8	2457451.646	2457451.643± 0.001	-0.003	Tarot - Calern
2457451	V1361 Cen	2	0.899	0.900	9.6	9.6				Tarot - Chile
2457451	BI Cen	20	0.511	0.728	0.3	5.5	2457451.558	2457451.568± 0.003	0.011	Tarot - Chile
2457451	ET Hya	22	0.545	0.783	1.1	6.8	2457451.650	2457451.652± 0.003	0.003	Tarot - Chile
2457451	IU Car	24	0.595	0.799	2.3	7.2	2457451.687	2457451.694± 0.004	0.006	Tarot - Chile
2457451	TY Aps	22	0.644	0.860	3.5	8.6	2457451.670	2457451.672± 0.003	0.002	Tarot - Chile
2457451	V476 Vir	34	0.727	0.899	5.4	9.6	2457451.827	2457451.894± 0.005	0.067	Tarot - Chile
2457451	V689 Car	16	0.501	0.706	0.0	4.9	2457451.573	2457451.569± 0.004	-0.005	Tarot - Chile
2457451	X Crt	34	0.714	0.891	5.1	9.4				Tarot - Chile

Routine followup of bright RR Lyr stars, type RRab

An example: Blazhko star **AF Vel** observed at La Silla over 10 years, 68 maxima, 2458 measurements







Human observers: Study of understudied stars

Many RR Lyr satrs fainter than magnitude 13 at minimum are understudied.

Pulsation period and type are not accurately known

We do not know if there is a Blazhko effect

Observers use telescopes of 20 to 60 cm diameter

Observing plan: get complete folded light curve in a short time (about one week) and do it again several weeks or months later.

Variable number of amateur astronomers contribute to the program Presently from Europe and Japan Since 10 years, 25 astronomers have contributed, from Europe, Japan, Australia, USA and Namibia

~100 unserstudied RR Lyr stars observed

Several Blazhko effect discovered

An example : CM Leo

Known as RRab proved to be an RRc

Observers: 2011-2015

Laurent Corp and collaborators, T60 Pic du Midi, France

Marco Nobile, 20cm Savosa, Switzerland Maurice Audejean, 30cm, Chinon, France

Also Francesco Fumagalli and Gisela Maintz (BAV)

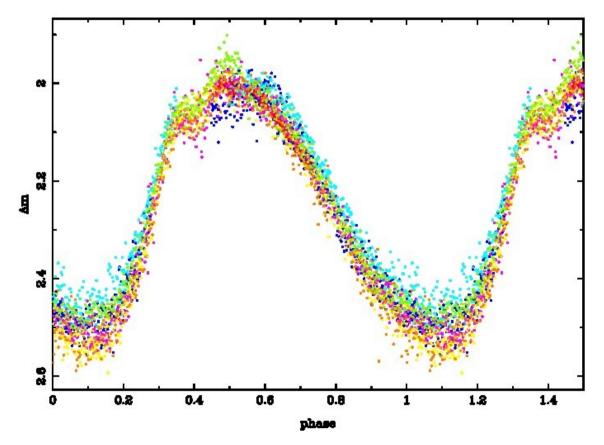
CM Leo 3357 mes.

Observed period: 0.366 jour

GCVS: RRAB

Magnitude: 13.8-14.9 (B)

Period: 0.361732 jour



V568 Cas : A long period Blazhko effect

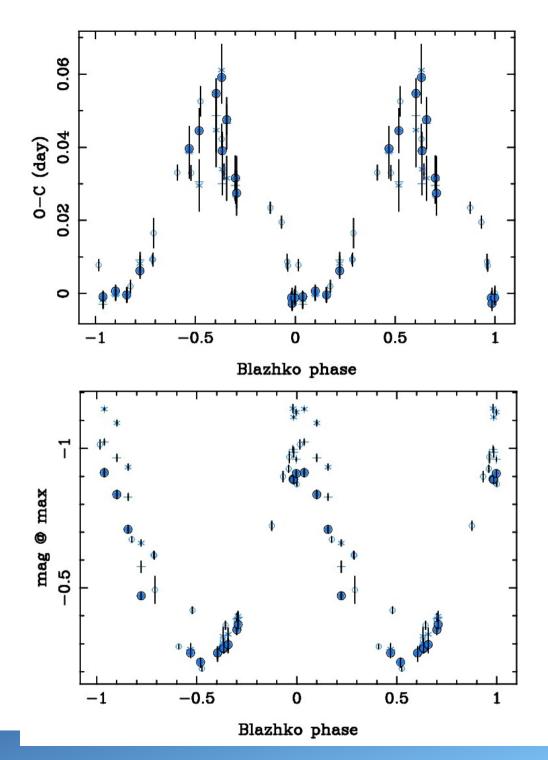
GCVS: RR:

Magnitude 13.1-14.0 (p)

Period: 0.623 jour

Observed period: 0.51404 days Blazhko effect period: 325 jours

Observers: 2008-2016 Eric Denoux, Caussade, France, 28cm Mercè Correa, Sabadell Spain, 50cm Ramon Moliner, Sabadell Spain, 50cm



Followup of the Blazhko effect of RR Lyr itself

RR Lyr magnitude 7-8, period 0.51 jour Blazhko effect period ~40 days.

Variable Blazhko effect:

Preston et al. (1965) observed disappearance of Blazhko effect in 1963.

The phenomenon has a time scale of a few years,

Study of Variable Blazhko effect needs continuous observation during several years,

Instrumentation : **VTT,** Design : Alain Klotz

Camera CCD Audine kaf400 Photo lens 135mm F/2.8 Mount HEQ5 GOTO controlled by computer

A job for ants!



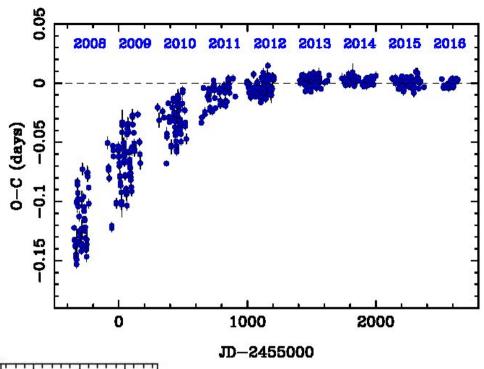


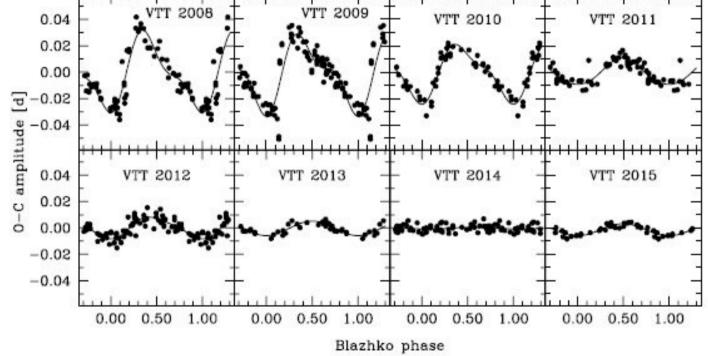
2008-2016:

2 VTTs:980 nights of observations394 maxima

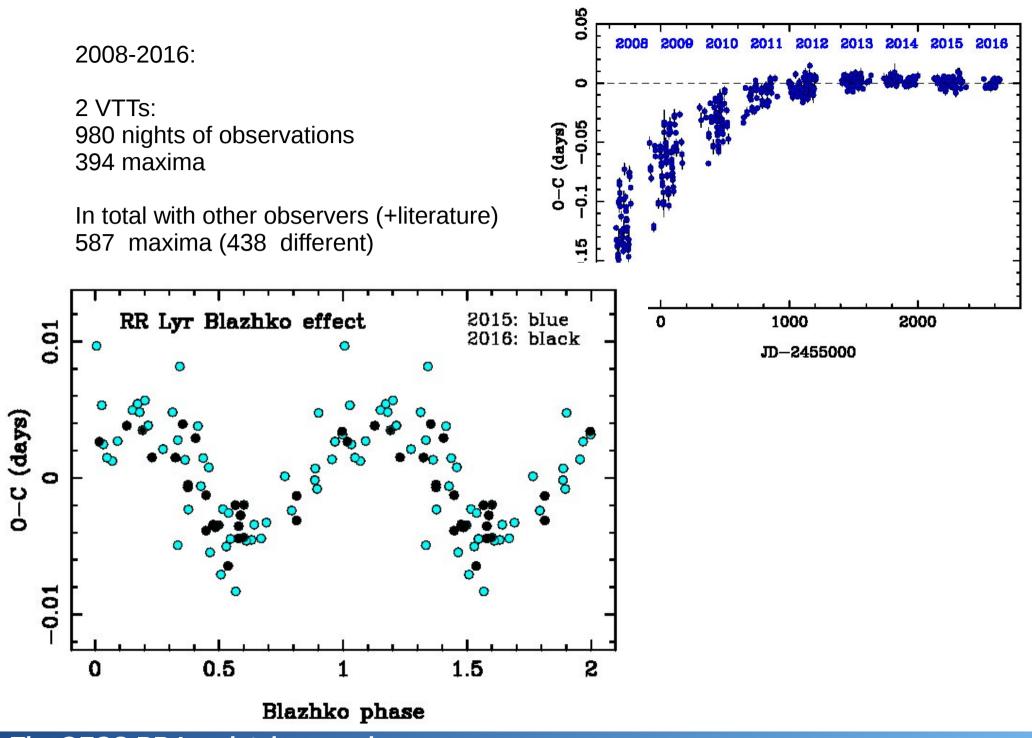
In total with other observers (+literature) 587 maxima (438 different)

Disappearance of Blazhko effect in 2014 as in 1963 (Le Borgne et al. 2014) and Restart in 2015





From Poretti, Le Borgne, Klotz, Audejean, Hirosawa, 2015, Conference "RRL2015 - High-Precision Studies of RR Lyrae stars", held in Visegrad (Hungary)



Conclusion

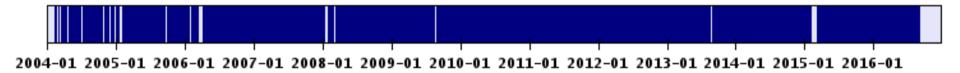
To find new things on RR Lyr stars, no need of giant telescopes.

But need to observe a lot (short time sampling, minutes), during a long time (years): this is slow science!

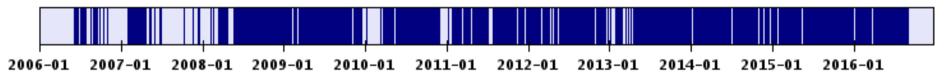
Solution: collaborative work and use of automated/robotic telescopes.

GRRS observation density 2004-2016:

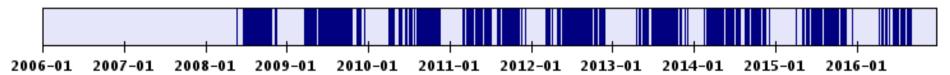
Robotic telescopes Tarot:



Humans, understudied stars:



VTTs, RR Lyr itself:



Thank you